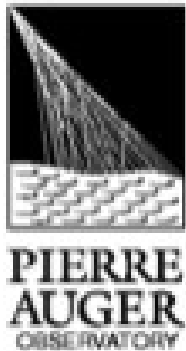


# Simulation of Extensive Air Showers for Auger on GRID



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Pierre Auger Collaboration  
University of Granada



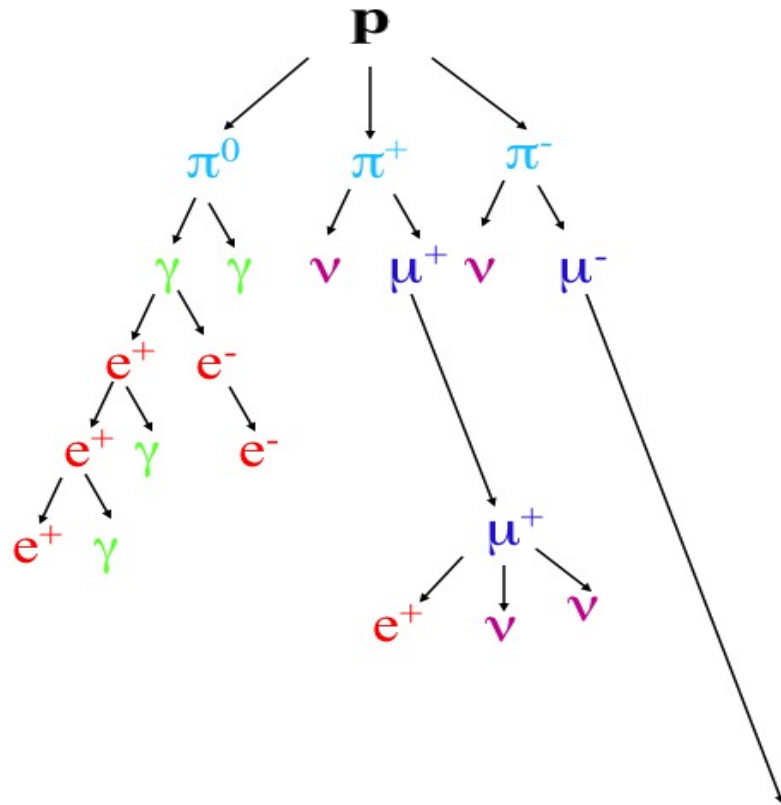
# Simulation of Extensive Air Showers

## Shower

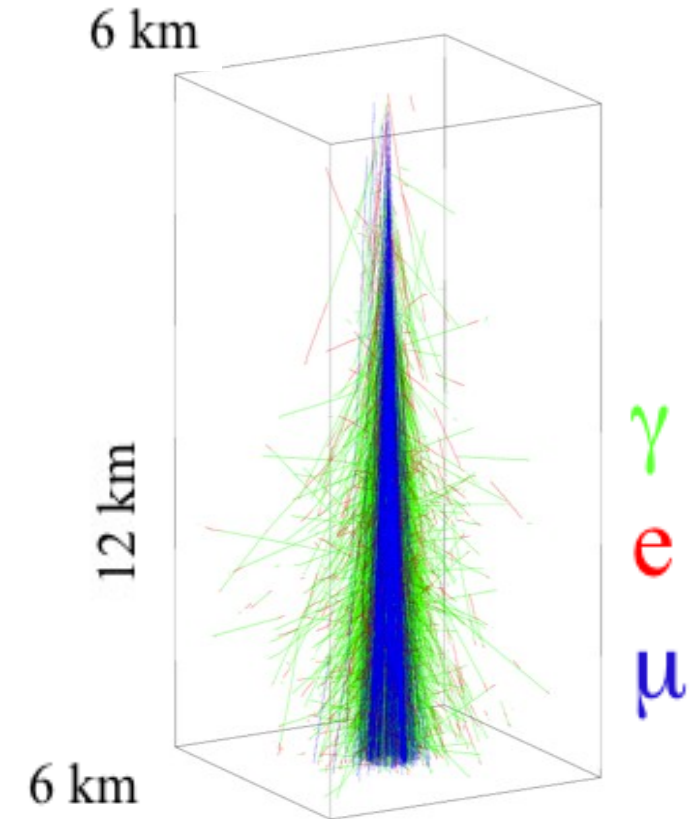
Cosmic ray primary interacts creating mostly secondary pions :

Neutral pions : electromagnetic shower

Charged pions : decay in muons



**$10^{19}$  eV proton**



Simulation of a  $10^{19}$  eV proton EAS using the MOCCA program. A sample of tracks at  $> 300$  m from the shower axis are shown. Frame box:  $6 \times 6 \times 12$  km high. Color code:  $\gamma$  green,  $e$  red,  $\mu$  blue. Drawn by Clem Prike — University of Chicago

**Simulation by Clem Prike**

# Simulation of EAS

## Shower generation:

- ∅ **CORSIKA** and **AIRES** are the software packages that generate those kind of events. In official simulations we have only used CORSIKA
- ∅ A compilation tool lets the user decide on:
  - ∅ **Low energy** interaction models
  - ∅ **High energy** interaction models (epos, QGSjetI/II, ...)
  - ∅ and some other options
- ∅ CORSIKA requires only an **input card** (run number, particle type, energy, zenith angle, seeds, etc ... ) which is specific for each job.
- ∅ Billions of particles being tracked:
  - ∅ A **full shower** simulation can consume **hundreds of hours of CPU time and several TBs of disk space**.
  - ∅ **Statistical thinning method** is needed, but **still tens of hours of CPU and hundreds of MB** of disk space are required for a single shower.

# Pierre Auger Observatory

It is located near Malargüe, south of Mendoza in Argentina

- ∅ **Red dots** : water tanks (SD)
  - ∅ **Yellow labels** : location of telescope eyes (FD)
- Green lines: coverage in azimuth



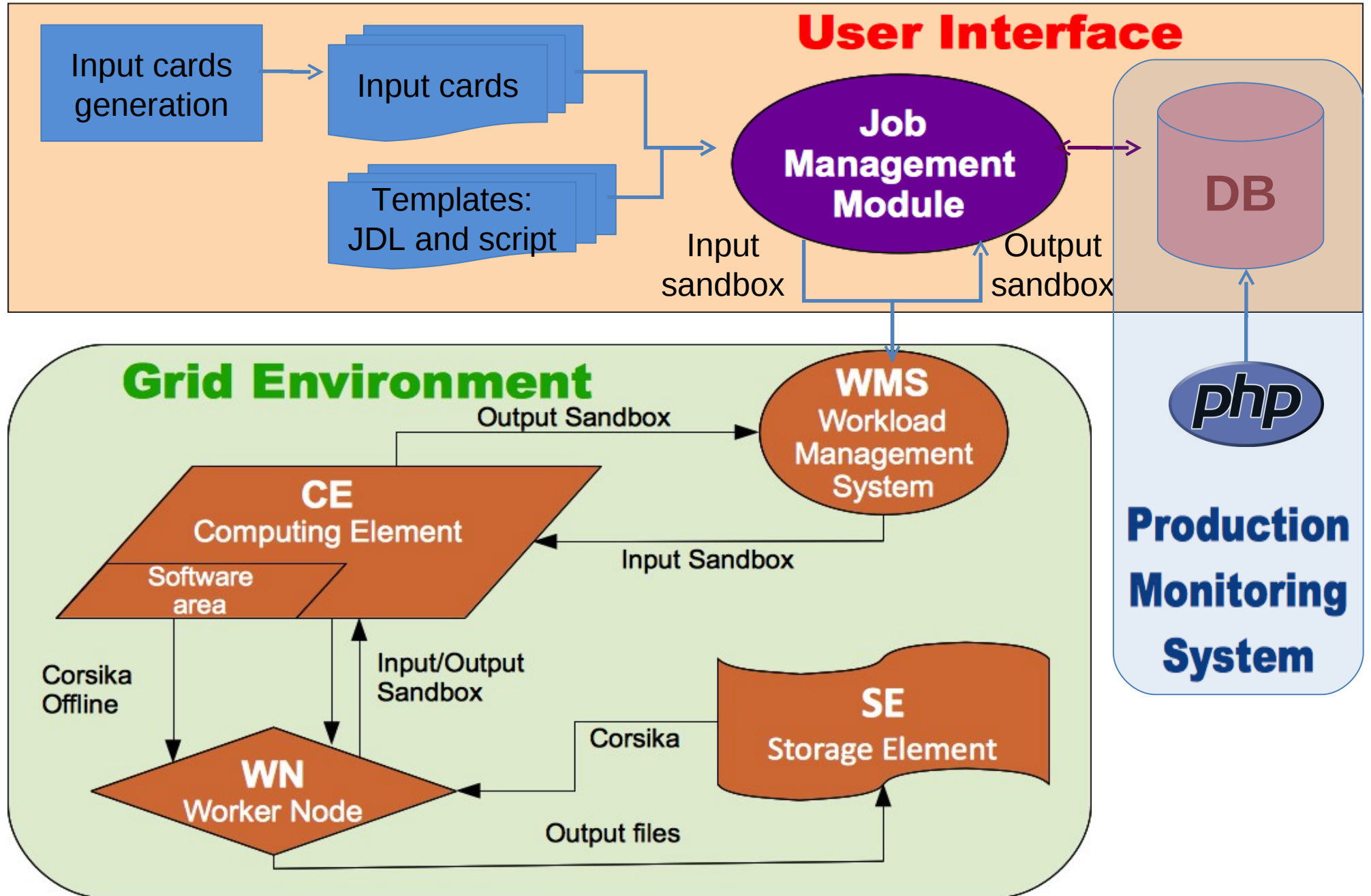
# Detector sim. and rec.

## Detector simulation and reconstruction:

- ∅ Needs previously generated shower files → showers have to be kept on Grid
- ∅ Auger experimental devices detect :
  - ∅ Cherenkov light of particles remaining at ground level as secondaries from the shower (SD)
  - ∅ Fluorescence light emitted while those secondaries are traveling through the atmosphere (FD)
  - ∅ In both cases the output of the detectors are FADC traces (not much data)
- ∅ **OffLine**: modular package to simulate the response of the SD and FD and to reconstruct physical parameters of the shower.
- ∅ File sizes are usually of the order of MBs

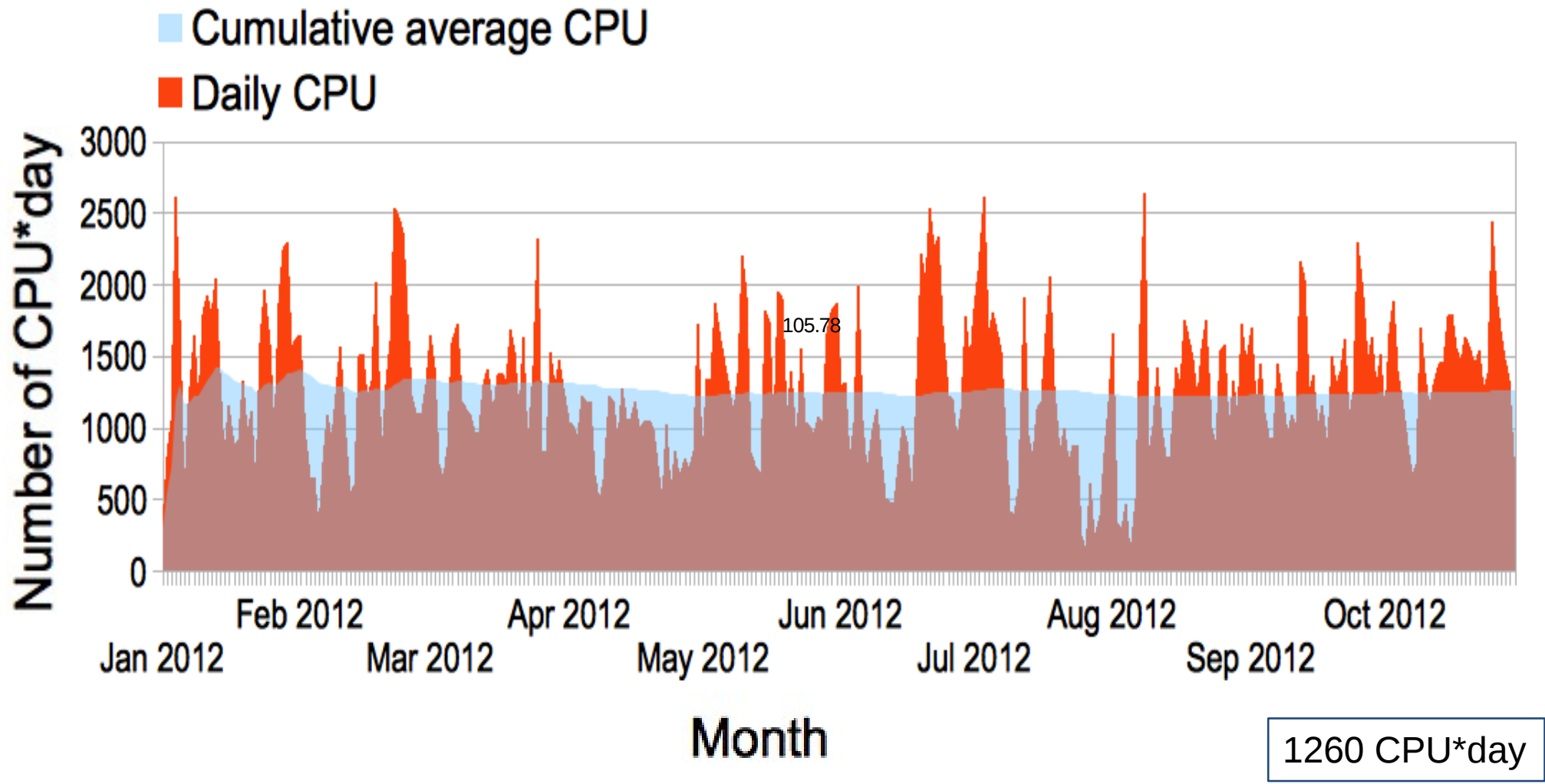


# Grid Computing Model



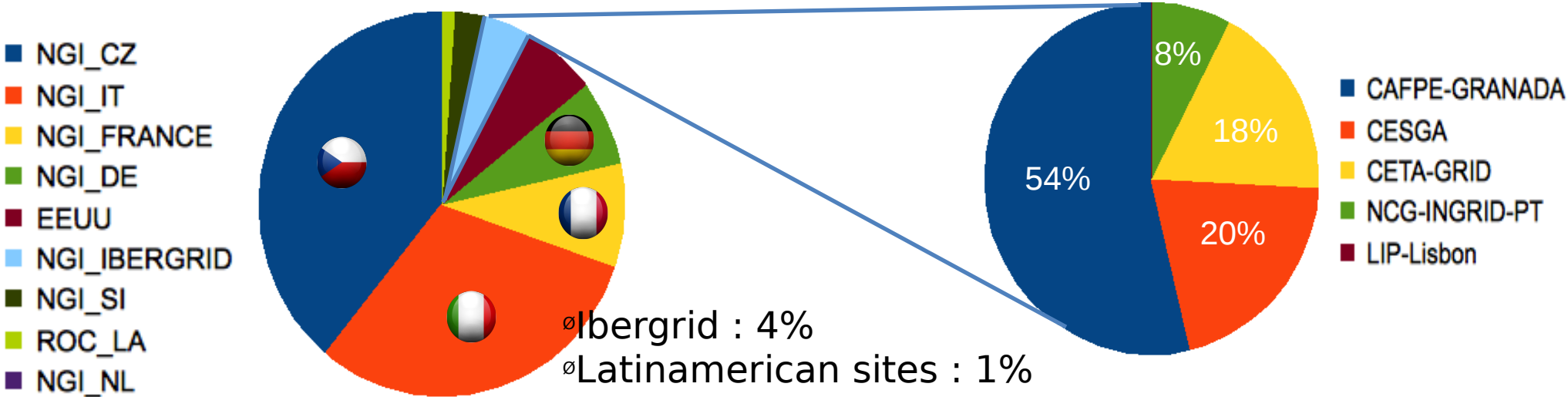
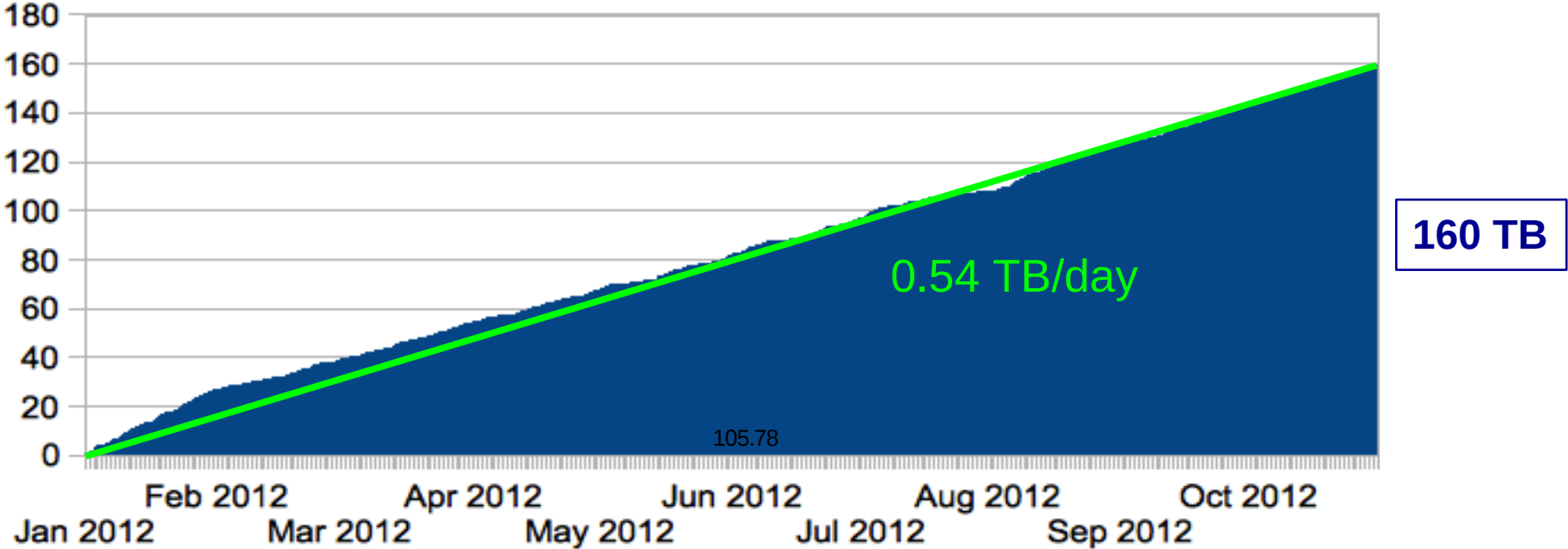
# Usage of resources 2012

Consumption of CPU time since the start of 2012



# Usage of resources 2012

## Disk space accumulated on Storage Elements

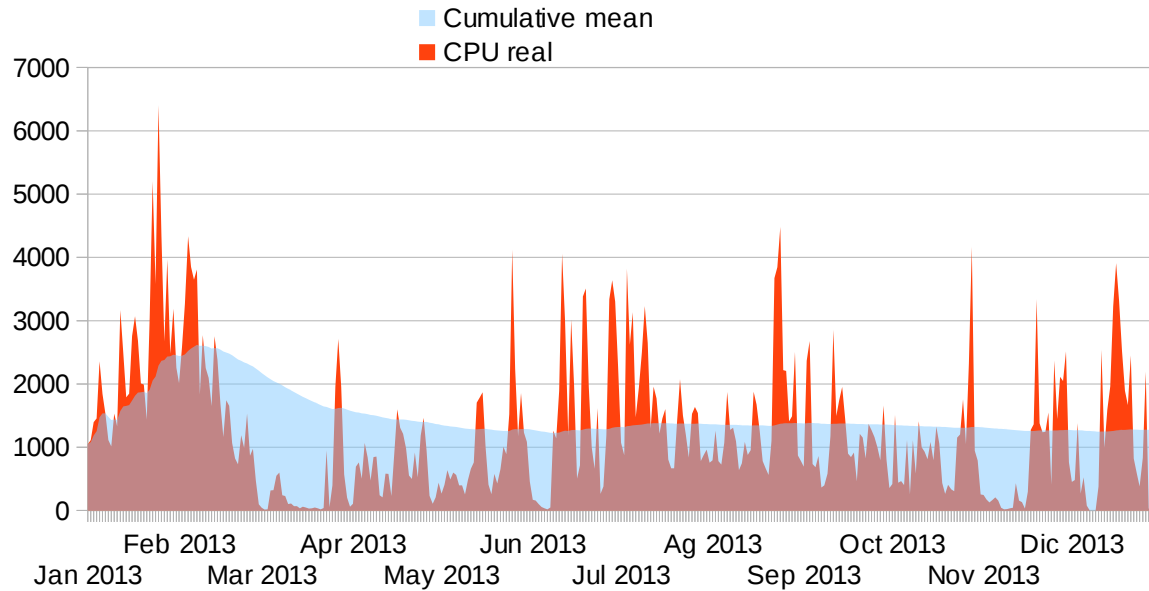




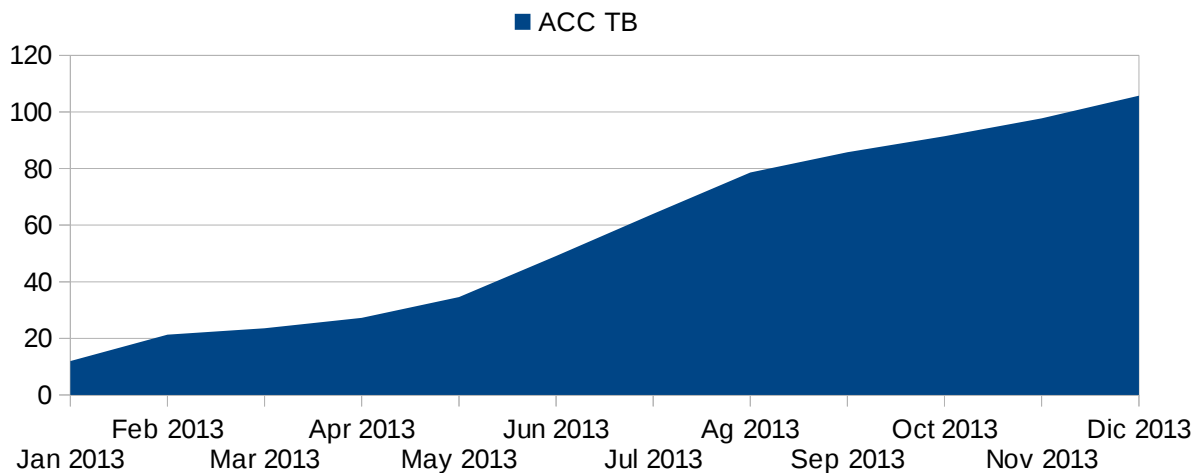
# Usage of resources 2013

## Consumption of CPU time since the start of 2013

Number of CPU\*day



1291 CPU\*day



106 TB